Boundary Terms in Gravitational Action

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- Action principle in general relativity.
- Constructing a well-posed action for general relativity as well as in Lovelock theories in relation to null surfaces.

Main References

- SC and Parattu, arXiv:1806.08823.
- SC, Parattu and Padmanabhan, arXiv:1703.00624.
- **SC**, arXiv:1607.05986.
- Parattu, **SC** and Padmanabhan, arXiv:1602.07546.
- Parattu, SC, Majhi and Padmanabhan, arXiv:1501.08765.

• The covariant action for general relativity is the Ricci scalar. In four dimension it yields,

GR Action

$$16\pi G\mathcal{A} = \int_{\mathcal{V}} d^4x \; \sqrt{-g} R(g, \partial g, \partial^2 g)$$

• The Ricci scalar is constructed from the Riemann tensor, which contains second derivatives of the metric.

Variation of the Action

- Vary the gravitational action.
- We will consider the difference of the action for two metric configuration, g_{ab} and a varied one, $g_{ab} + \delta g_{ab}$.
- This leads to,

variation

$$16\pi G \delta \mathcal{A} = \int_{\mathcal{V}} d^4 x \sqrt{-g} \left(R_{ab} - \frac{1}{2} Rg_{ab} \right) \delta g^{ab} \\ + \epsilon \int_{\partial \mathcal{V}} d^3 y \sqrt{|h|} n^a h^{bc} \left(\partial_b \delta g_{ac} - \partial_a \delta g_{bc} \right) \\ + \text{Boundary Terms with } \delta g^{ab}$$

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A possibility to make the action well-posed SC, Parattu and Padmanabhan, arXiv:1703.00624 Charap and Nelson, J. Phys. A 16, 1661 (1983)

• There are many alternatives to make the Einstein-Hilbert action well-posed. Among them the boundary term proposed by Gibbons and Hawking is the famous one.

Basic Idea

$$\delta \mathcal{A} = \int d^4 x$$
 (Equation of Motion Term) δ (Dynamical Variable)
+ $\int d^3 x$ (Conjugate Momentum) δ (Variables to be fixed)
+ $\int d^3 x \, \delta$ (Boundary Term) + $\int d^3 x$ (Total Divergence Term)

Is it Possible? Padmanabhan, MPLA 29, 1450037 (2014) Gibbons and Hawking, Phys. Rev. D 15, 2752 (1977) York, Phys. Rev. Lett 28, 1082 (1972)

• The variation of the Einstein-Hilbert action involves,

Variation on Surface

$$\mathcal{B} = \text{Tot. Derv.} - \delta \left(2 \nabla_a n^a \right) + \left(\nabla_i n_j - \epsilon n_i a_j \right) \delta g^{ij}$$

• Using the properties of the extrinsic curvature

Finally

$$\sqrt{h}\mathcal{B} = \text{Tot. Derv.} - \delta\left(2K\sqrt{h}\right) + (Kh_{ij} - K_{ij})\delta h^{ij}\sqrt{h}$$

• $2K\sqrt{h}$ term to be added to the gravitational action.

Surface Variation on a null surface

Parattu, SC, Majhi and Padmanabhan, arXiv:1501.08765

• Starting from the surface variation we can evaluate it for a null surface as well. This will lead to,

Surface Variation

$$\mathcal{B} = \text{Tot. Derv.} - 2\delta \left[\sqrt{q} \left(\Theta + \kappa \right) \right] + \sqrt{q} \left[\Theta_{ab} - \left(\Theta + \kappa \right) q_{ab} \right] \delta q^{ab} + 2\sqrt{q} \left[\left(\Theta + \kappa \right) k_c + \ell^b \nabla_c k_b \right] \delta \ell^c$$

• Thus for null surfaces we need to add a term $2\sqrt{q} (\Theta + \kappa)$ to the Ricci scalar.

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• When this modified action is varied we will get two surface contributions, one from δq^{ab} and another from $\delta \ell^c$.

- Can we derive the structure of the boundary term in the form language? What happens when the boundaries are not smooth? These have been answered in:
 - I. Jubb, J. Samuel, R. Sorkin and S. Surya, CQG 34, 065006 (2017).
 - 2 L. Lehner, R.C. Myers, E. Poisson, R.D. Sorkin, PRD 94, 084046 (2016) and others.
- The proposal "Complexity of a boundary state in CFT = Gravitational Action in the bulk" crucially requires the boundary term for null surfaces. These results appear in:
 - S. Chapman, H. Marrochio and R.C. Myers, JHEP, 1701, 062 (2017)
 - J. Maltz and L. Susskind, Phys. Rev. Lett. 118, 101602 (2017) and others.
- These results have also been used in various other contexts, e.g., de Sitter spacetime, degrees of freedom, AdS Soliton etc.

Variation of the Gauss-Bonnet Lagrangian

SC, Parattu and Padmanabhan, arXiv:1703.00624

Myers, Phys. Rev. D 36, 392 (1987)

Variation

Surface Variation = Total Divergence + (Conjugate Momentum) δ (Variables to be fixed) + $\delta \left\{ 8 {}^{(D-1)} R_b^a K_a^b - 4 {}^{(D-1)} R K \right\}$ + $\epsilon \left(\frac{4}{3} K^3 + \frac{8}{3} K_b^a K_c^b K_a^c - 4 K K_b^a K_a^b \right) \right\}$

• A similar calculation yields the corresponding boundary term for Lanczos-Lovelock gravity as well.

Null Boundary Terms in Gauss-Bonnet Gravity SC and Parattu, arXiv:1806.08823

Boundary Term Gauss-Bonnet

$$\begin{split} \sqrt{q} \ \mathcal{B}_{\rm GB} = & 4\sqrt{q} \bigg[\kappa \left(\begin{array}{c} (D-2)R - 2\Theta_q^p \Psi_p^q + 2\Theta \Psi \right) + 2\Psi_q^p \mathcal{L}_\ell \Theta_p^q \\ & - 2\Psi \frac{d\Theta}{d\lambda} + 2\Theta_c^a \Theta_b^c \Psi_a^b - 2\Psi \Theta_b^a \Theta_a^b \\ & - 4 \left(D^m \Theta - D^a \Theta_a^m \right) \left(k^q \nabla_m \ell_q \right) \\ & - 2 \left(k^q \nabla_m \ell_q \right) \left(k^p \nabla_a \ell_p \right) \left(\Theta q^{ma} - \Theta^{ma} \right) \bigg] \end{split}$$

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- Elaborated on the associated complications with gravitational action principle.
- Possible resolutions to make the variational principle well posed for null surfaces.
- The correct action associated with spacelike/timelike and null surfaces in the context of Lanczos-Lovelock gravity have been put forward.

Thank You

Sumanta Chakraborty Null Surfaces

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